

Photoemission Process

Photo emission process

$$\Psi_{\text{tot}}^i(N), E_{\text{tot}}^i(N) \rightarrow \Psi_{\text{tot}}^f(N, K), E_{\text{tot}}^f(N, K)$$

$\Psi_{\text{tot}}^f(N, K) \rightarrow K^{\text{th}}$ final state of N electron system.

(involves both the emitter and the photoelectron)

Energy conservation equation.

$$E_{\text{tot}}^i(N) + h\nu = E_{\text{tot}}^f(N, K)$$

k refers to one electron orbital k from which electron emission occurred, in the simplest case. In general it should include all modes of excitation, electronic, vibrational and translational.

In general, it is assumed that the photoelectron is weakly coupled to the (N – 1) electron system so that separation of final state is possible

The binding energy corresponding to producing the $\Psi_{\text{tot}}^f(N-1, K)$ state is,

$$E_b^v(K) = E_{\text{tot}}^{(f)}(N - 1, K) - E_{\text{tot}}^i(N)$$

The photoelectron is referred w.r.t the vacuum level.

Line width \rightarrow lifetime of initial and final states

Initial lifetimes are large, for a final state life time t the lorentzian width of the peak is $\sim \hbar/t = 6.58 \times 10^{-16}/t(\text{s})$ in eV. For life time of 10^{-18} sec this width can be large (initial state lifetime in most cases is high and there is no contribution to line width)

System containing N electrons with spatial coordinates $r_1, r_2 \dots r_N$ and spin coordinates $\sigma_1, \sigma_2, \dots \sigma_N$ and p nuclei with coordination $R_1, R_2 \dots R_N$ the total wave function,

$$\Psi_{\text{tot}}(N) = \Psi_{\text{tot}}(r_1, \sigma_1, r_2, \sigma_2 \dots r_N \sigma_N; R_1, R_2 \dots R_P)$$

Nuclear coordinates can be neglected in the resolution scale of electron spectroscopy. In the non-relativistic limit, the Hamiltonian in electrostatic units,

$$\hat{H}_{\text{tot}} = \underbrace{\hbar^2/2m \sum_{i=1}^N \nabla_i^2}_{\text{e-kinetic}} - \underbrace{\sum_{i=1}^N \sum_{l=1}^P Z_l e^2/r_{il}}_{\text{e-n attraction}} + \underbrace{\sum_{i=1}^N \sum_{j>i}^N e^2/r_{ij}}_{\text{e-e repulsion}} \\ + \underbrace{\sum_{l=1}^P \sum_{m>l}^P Z_l Z_m e^2/r_{lm}}_{\text{n-n repulsion}} - \underbrace{\hbar^2/2 \sum_{l=1}^P \nabla_l^2 / M_l}_{\text{Nuclear kinetic}}$$

M – electron mass, Z_l charge of l^{th} nucleus $r_{il} = |r_i - R_l|$, $r_{ij} = |r_j - r_i|$,
 $R_{lm} = |R_l - R_m|$, M_l – mass of l^{th} nucleus.

To this relativistic effects may be added by perturbation. Hamiltonian most often added is due to spin-orbit splitting.